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3 In order to obtain the field equations of $f(T,f)$ gravity to study its cosmological applications, we consider the homogeneous and isotropic flat Friedmann-Lemaître-Robertson-Walker (FLRW) space time as $ds^2 = dt^2 + a^2(t)[dx^2 + dy^2 + dz^2]$, (6) where $a(t)$ is the scale factor that represents the expansion rate in the spatial directions and the tetrad, eA^m